

# Gaia: Review, Status & News

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SFM, El Escorial, June 16, 2017

# Outline



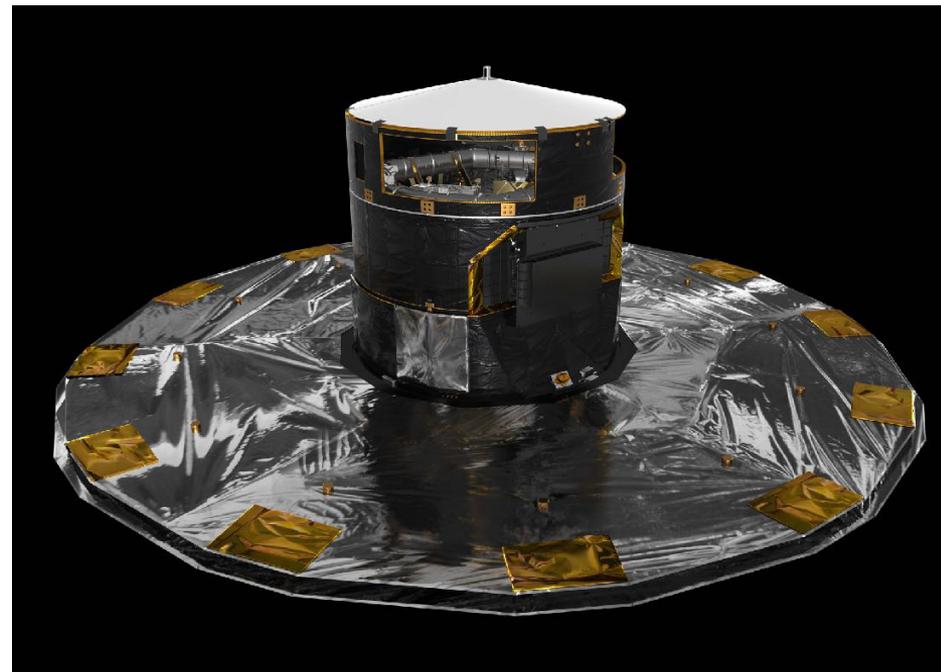
- Gaia, mission concept, operations, status
- Gaia DR1
- Plans for Gaia DR2
- Questions and Answers
- Thanks to Anthony Brown, Timo Prusti & Lennart Lindegren
- IAUS 330: Astrometry and Astrophysics in the Gaia Sky



# Gaia Summary

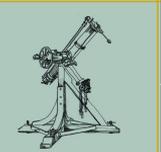
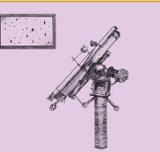
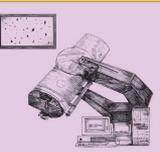
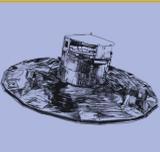


- ESA cornerstone mission building on the Hipparcos heritage
- Astrometry, Photometry and Spectroscopy
- Satellite, including the payload, by industry (Airbus DS), management and operations by ESA and data processing by scientists (DPAC)
- Launch 19 December 2013 with Soyuz from Kourou
- 5 years of operations in L2
- Gaia DR1 14 September 2016
- Science Alerts started 2014
- Gaia DR2 April-2018



# → ASTROMETRY THROUGH THE AGES



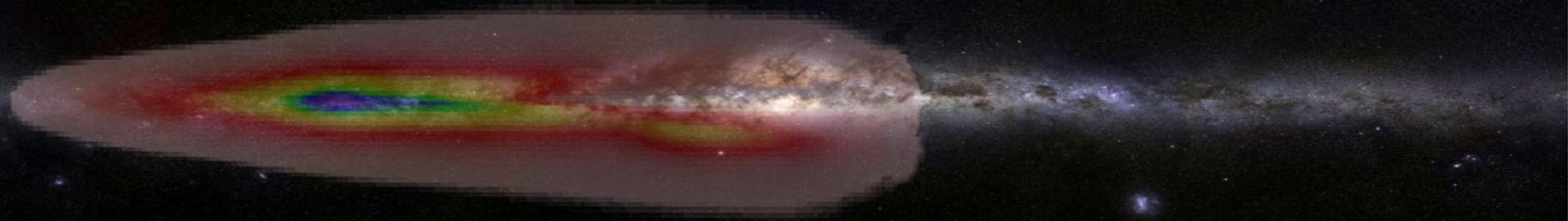
<p>* Hipparchus * — II century BCE —</p> 	<p>* Ulugh Beg * — 1437 —</p> 	<p>* Tycho Brahe * — 1598 (1627) —</p> 	<p>* John Flamsteed * — 1725 —</p> 	<p>* Jérôme Lalande * — 1801 —</p> 
<p>* Friedrich Bessel * Otto Struve * * Thomas Henderson * — 1837–1840 —</p> 	<p>* Jacobus Kapteyn * — 1910 —</p> 	<p>* Frank Schlesinger * * Louise Freeland Jenkins * * William van Altena * — 1924 — 1952 — 1995 —</p> 	<p>* Hipparcos * — 1989–1993 (1997) —</p> 	<p>* Gaia * — launched 2013 —</p> 

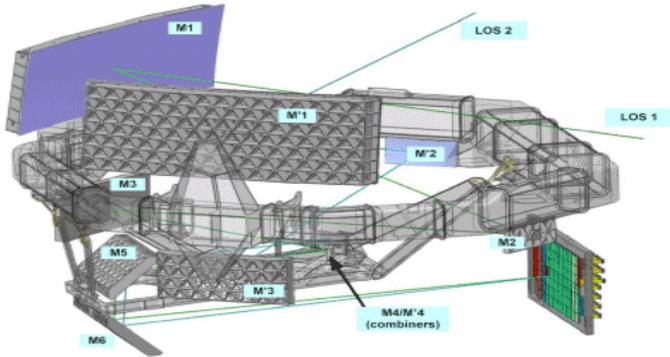
# The required reach of Gaia

'Our Sun'



# The required reach of Gaia



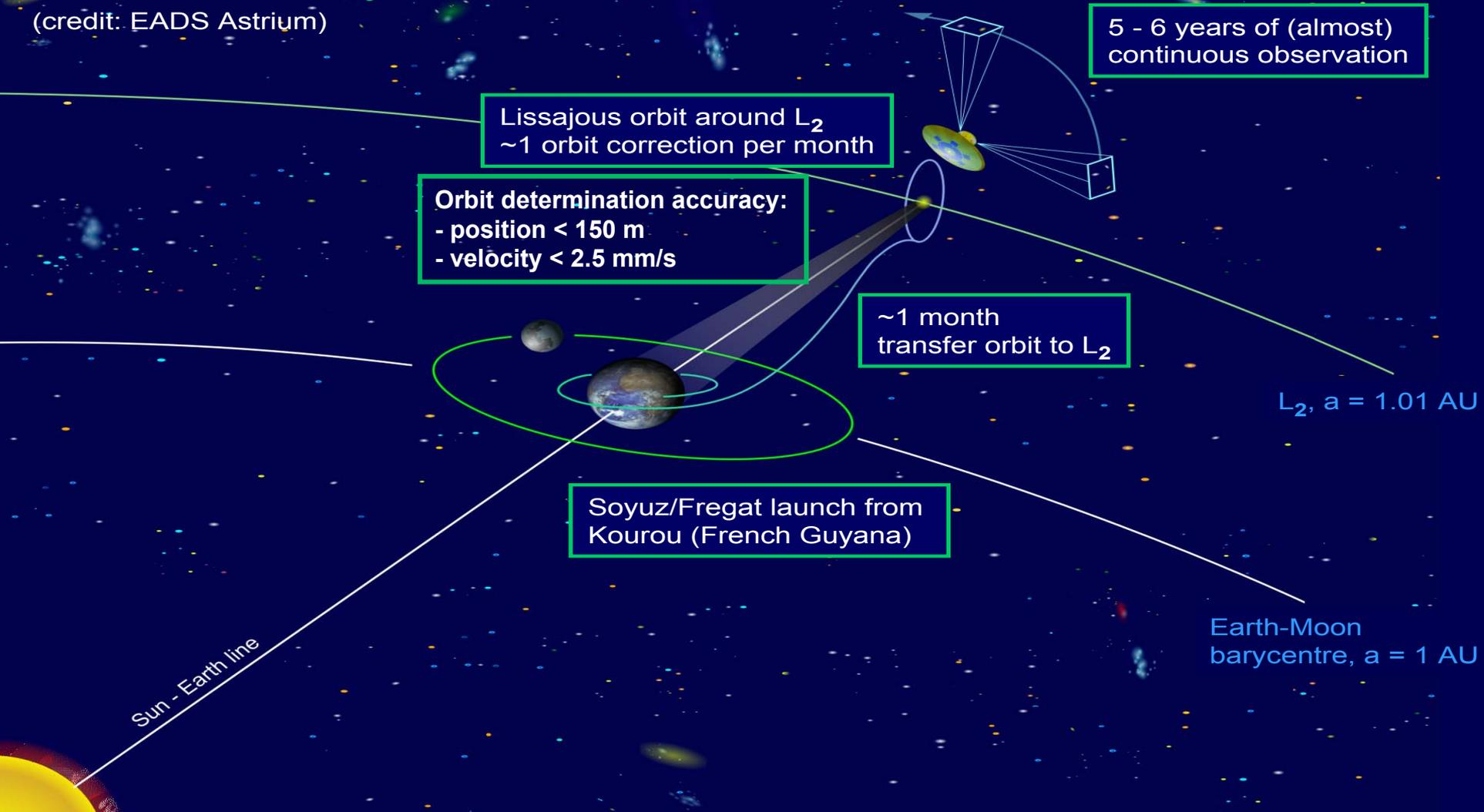


- How you can measure micro-arcsec:
  - Very stable thermal environment
  - New materials
  - Very Large Focal plane
  - Ultra-precise atomic clock
  - 80 observations of every source during 5 years
  - Very complex Data Processing, self calibration



# Gaia launch and orbit

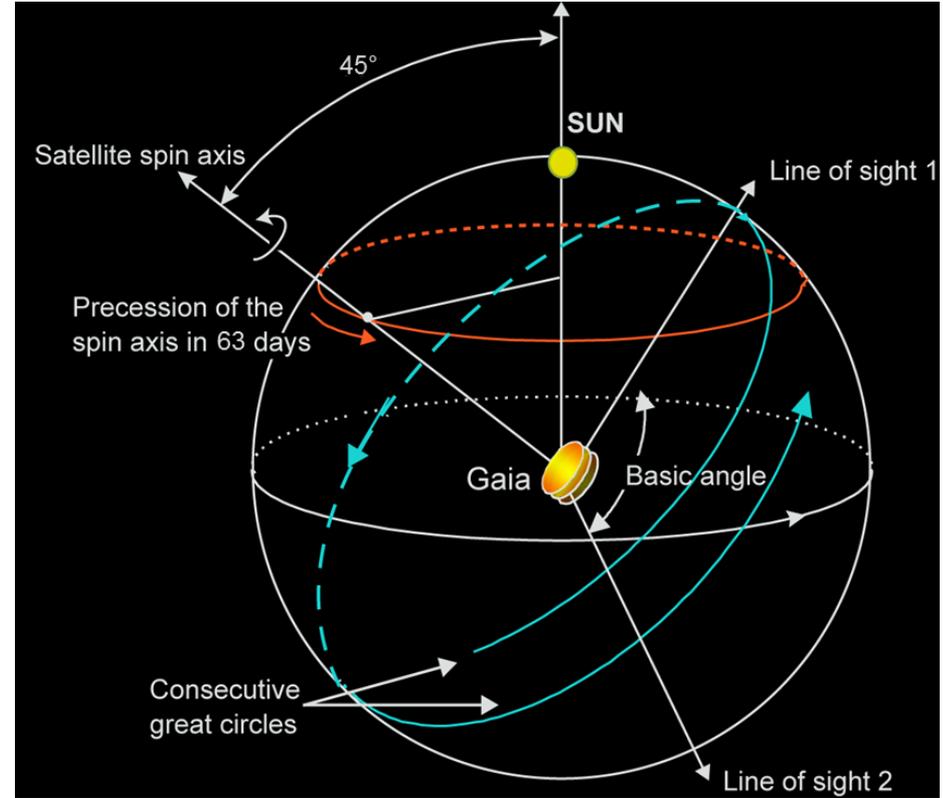
(credit: EADS Astrium)



# Gaia Operations

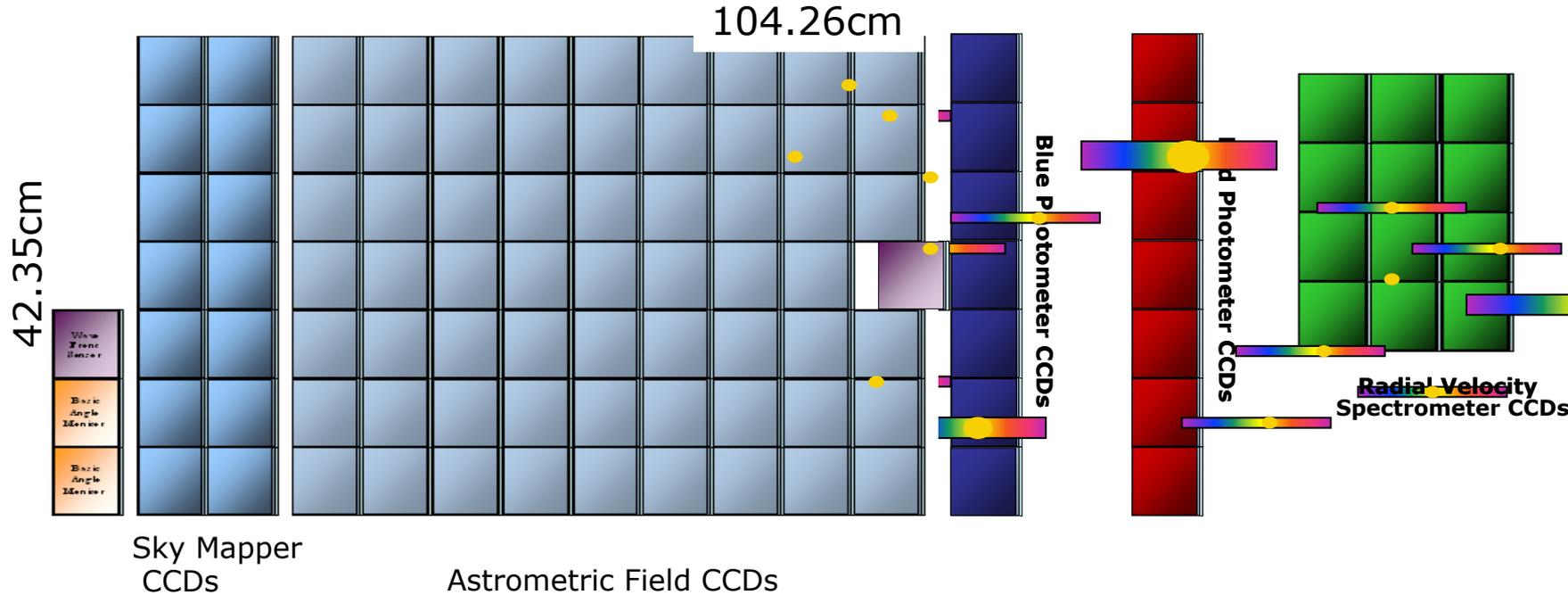


- Gaia in routine operations since July 2014
- 1,000 days routine phase reached 20 April
- Operations: nominal
- 76 billion transits observed
- Nominal 5-year mission ends mid-2019
- Estimated end of mission due to cold gas exhaustion end-2023 ( $\pm 1$  year)
- Mission extension process started



# One GigaPixel Focal Plane

**106 CCDs , 938 million pixels<sub>2</sub>, 2800 cm<sup>2</sup>**



- Not all the CCD data gets read and sent to Earth
- On board Video Processing Units read the windows around the detected stars
- Data is binned and stored compressed on board
- Data sent to Earth during the night ( $\sim 50$  Gb/day, 100Tb in 5 years)
- Three ESA DSN antennas used:
  - New Norcia
  - Cebreros
  - Malargüe
- Time shared with other missions

# Astrometric Measurements

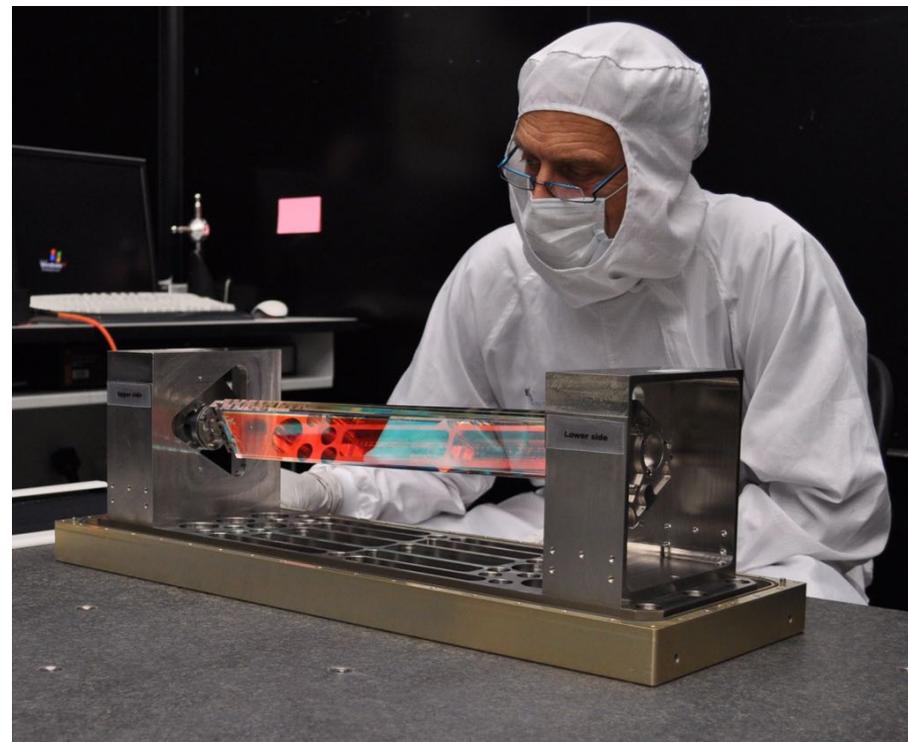


- Astrometric measurements: 761billion  $G < 20.7$  mag
- Bright limit around  $G = 2-3$  mag
- All bright stars covered with special measurements
- Selected crowded regions imaged with Gaia Sky Mapper



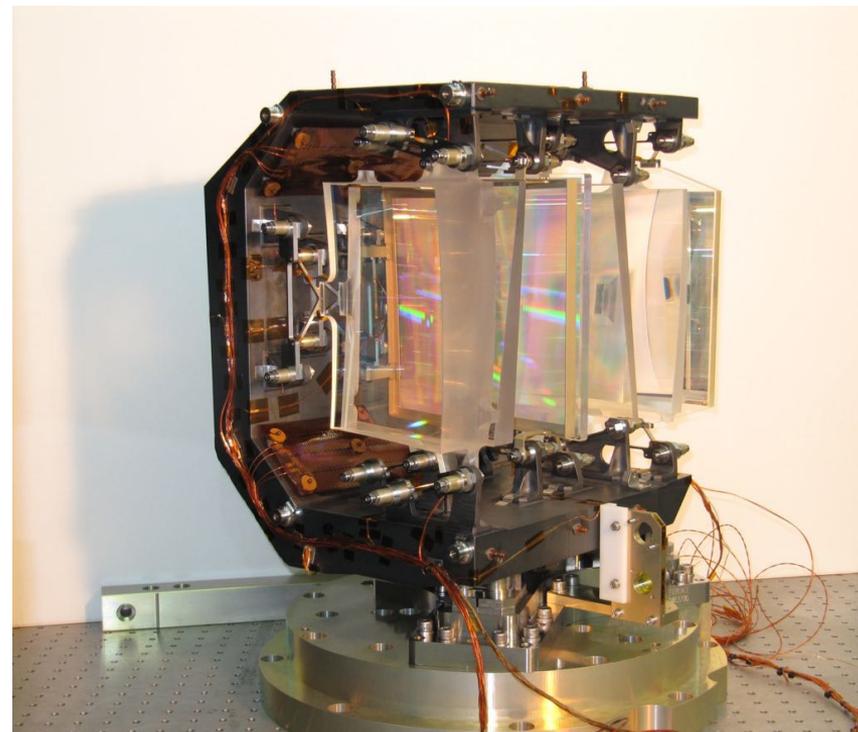
# Gaia Photometry

- Photometric measurements: 156 billion  $G < 20.7$  mag
  - Spectrophotometry
    - 330-680 nm BP
    - 640-1050 nm RP
  - Can also integrate BP and RP to get high precision measurements and a colour
- Astrometric measurements also photometric in G-band



# Gaia Spectroscopy

- Spectroscopic measurements: 13.7 billion
- GRVS < 16.2 mag
- 845-872 nm with R about 10,800
- Radial Velocity Spectrometer for >100 million radial velocities
- Spectroscopy till about GRVS=11-14 mag



# Unwanted features



- Contamination
  - Last decontamination in August 2016; no sign of transmission loss yet
- Micro-clanks and micrometeoroids
  - Taken into account in data processing for Gaia DR2
- Basic Angle Variation
  - Corrected with Basic Angle Monitor data for Gaia DR1 and DR2; more sophisticated analysis planned for the future
- Stray light
  - Impact on faint sources; on-board software modified from read-out dominated to background dominated case for faint objects
- Radiation damage
  - First signs visible, but less than anticipated before launch; pre-launch calibration work will become relevant in the future



# Gaia DR1

Credits: ESA/Gaia/DPAC  
Image acknowledgement: Moitinho & Barros  
Video acknowledgement: de Bruijne



NGC 6101  
NGC 6362

Mel 66

IC 4499

Carina

LMC NGC 2257

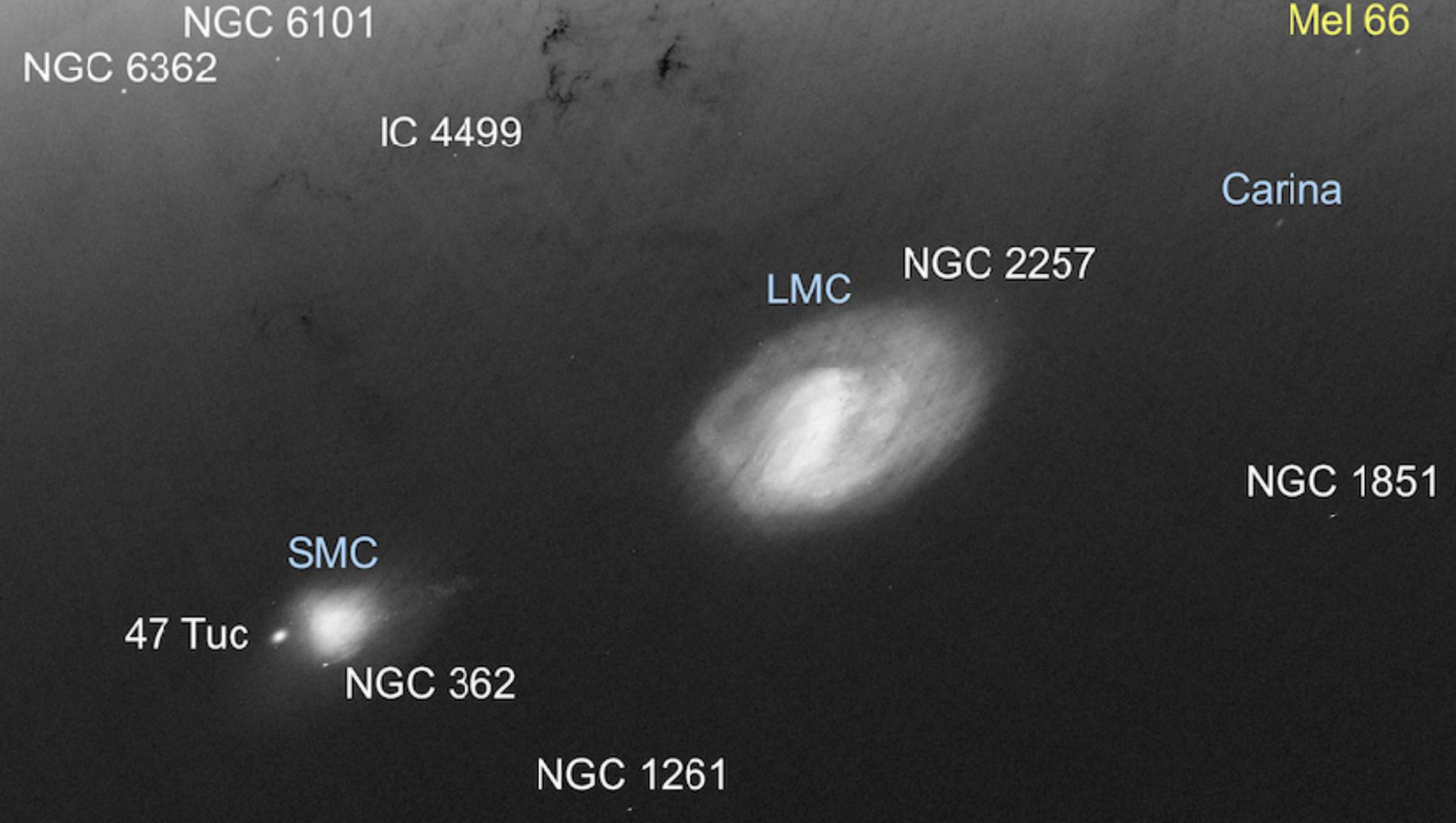
NGC 1851

SMC

47 Tuc

NGC 362

NGC 1261



NGC 2141

This is a grayscale astronomical image showing a field of stars. The stars are concentrated in a curved band across the frame. Several specific clusters are labeled with yellow text. The background is dark with some faint, diffuse light patterns.

NGC 2204

2243

M42

# Teamwork to deliver the promise of Gaia



- 10+ years of effort
- 450 scientists and engineers
- 160 institutes
- 24 countries and ESA
- Six data processing centres



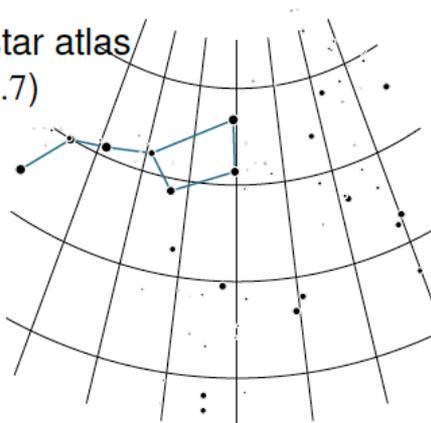
$\alpha \delta \varpi \mu_{\alpha^*} \mu_{\delta}$

IAUS330 - 2017.04.24 - 2/22

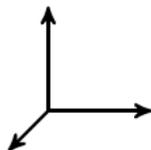


# What's in the Gaia DR1 delivery

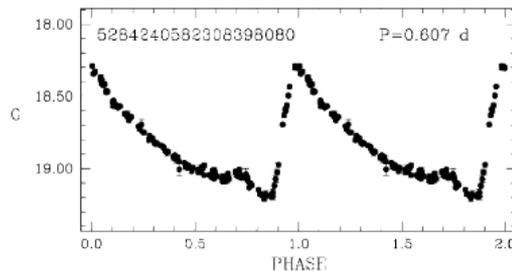
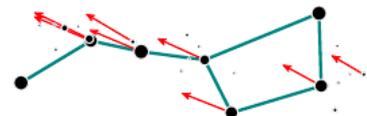
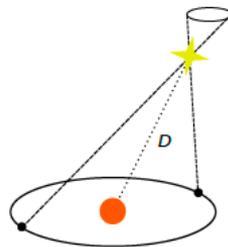
Billion star atlas  
( $G \lesssim 20.7$ )



Positions and magnitudes  
for  $\sim 2000$  ICRF quasars

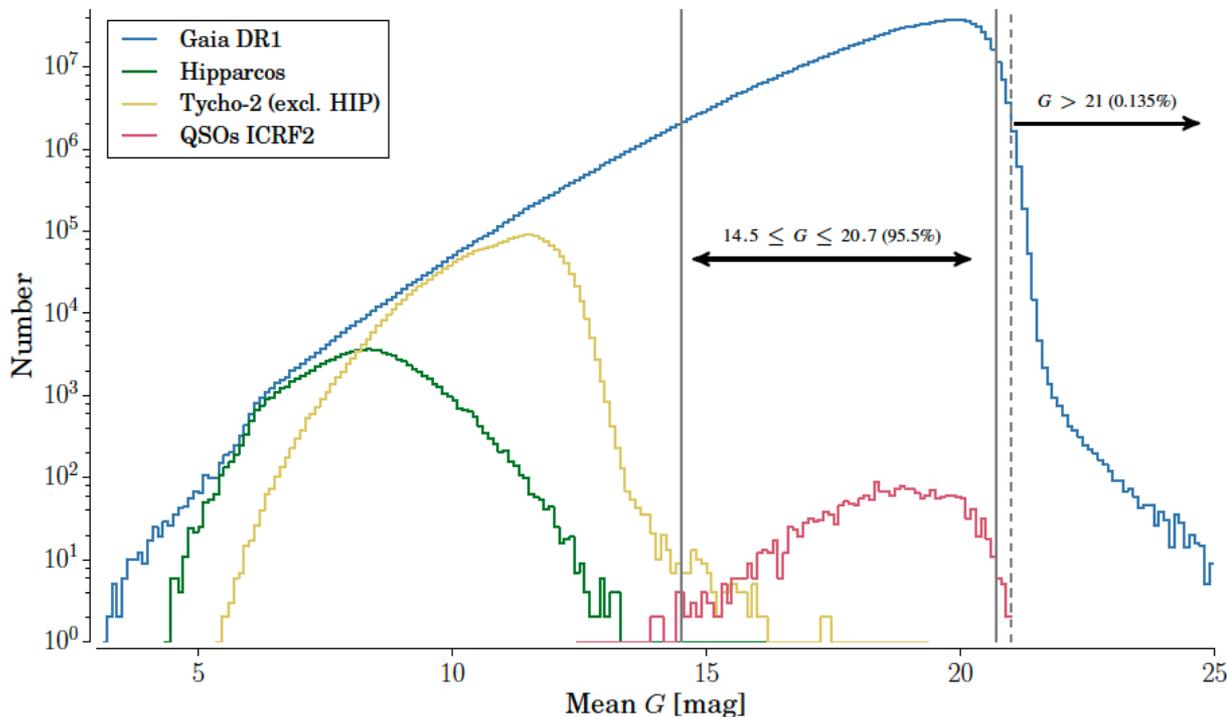


Tycho-Gaia  
Astrometric Solution  
( $\sim 2$  million,  $G \lesssim 12$ )



Variable stars near  
south ecliptic pole  
( $\sim 600$  Cepheids,  
 $\sim 2600$  RR Lyrae)

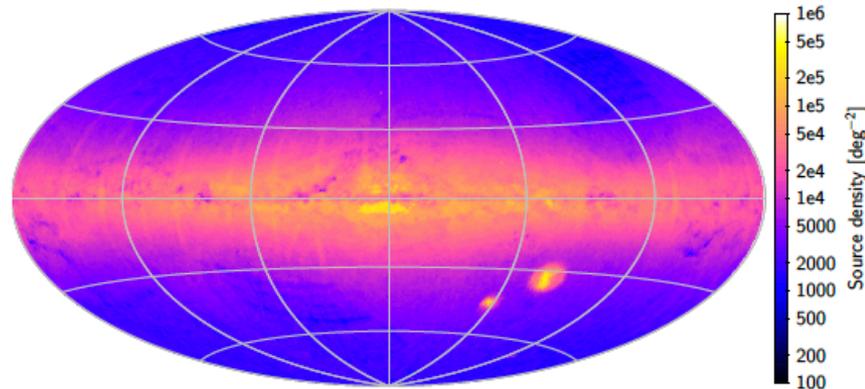
# Gaia DR1 magnitude distribution



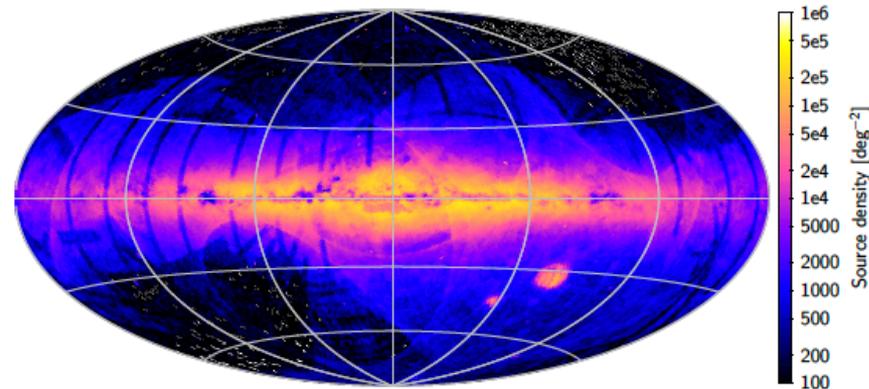
Gaia Collaboration, 2016, A&A

# Highly precise positions, new sources

685 million sources matched to IGSL



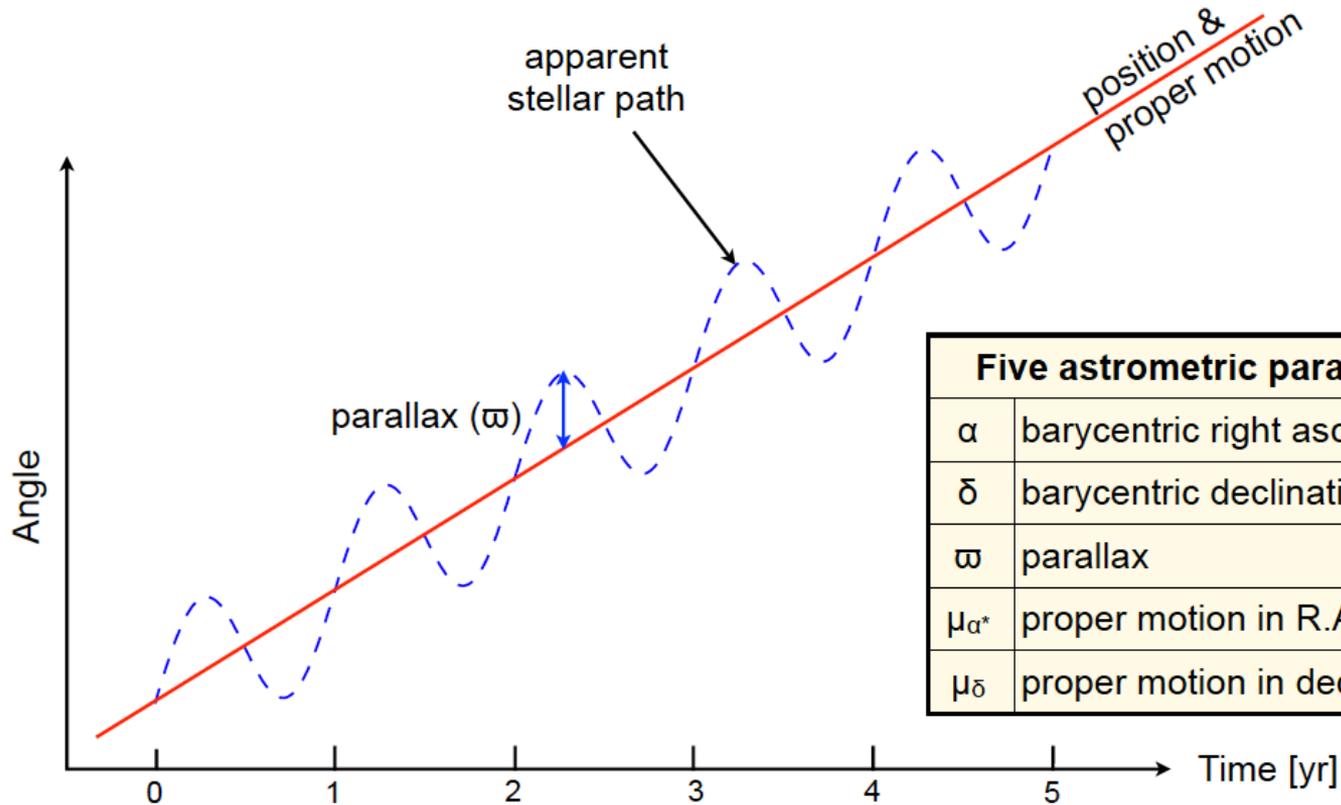
456 million new sources in Gaia DR1



DPAC/CU3/Lindgren et al., 2016, A&A

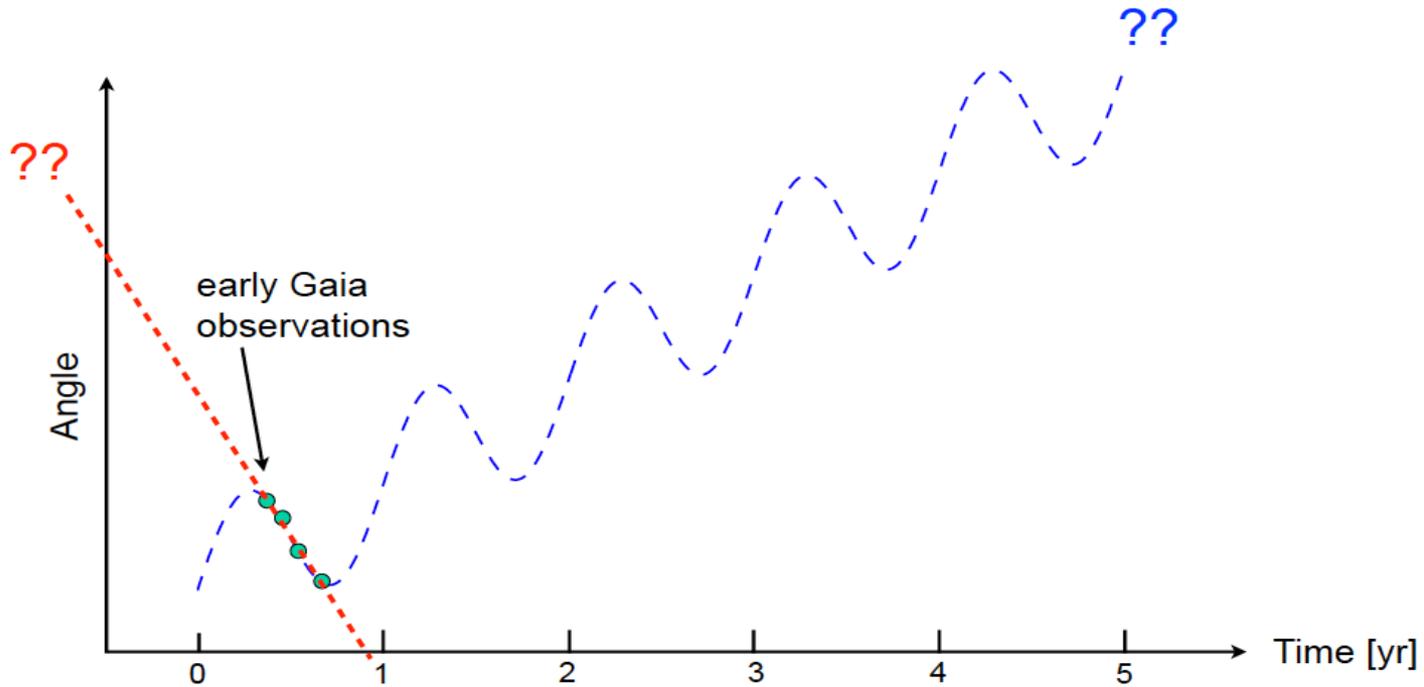
- $(\alpha, \delta)$  for  $\sim 1.1$  billion sources to  $G = 20.7$
- Epoch J2015.0, alignment to ICRF  $< 0.1$  mas, rotation  $< 0.03$  mas yr $^{-1}$
- Typical position uncertainty  $\sim 10$  mas
- Positions of 2191 ICRF sources from special astrometric solution (Mignard et al., 2016, A&A)
  - ▶ 90% with  $\sigma_{\text{pos}} < 3.35$  mas
  - ▶ no systematic differences with radio positions of more than few tenths of mas

# Astrometric model for a single star



Five astrometric parameters per star	
$\alpha$	barycentric right ascension at 2015.0
$\delta$	barycentric declination at 2015.0
$\varpi$	parallax
$\mu_{\alpha^*}$	proper motion in R.A. ( $\times \cos \delta$ )
$\mu_{\delta}$	proper motion in declination

# Degeneracy for <1 yr of observations



# TGAS: Using the old position as prior

Only **positions** were taken from Hipparcos and Tycho-2, no parallaxes or proper motions!

early Gaia observations

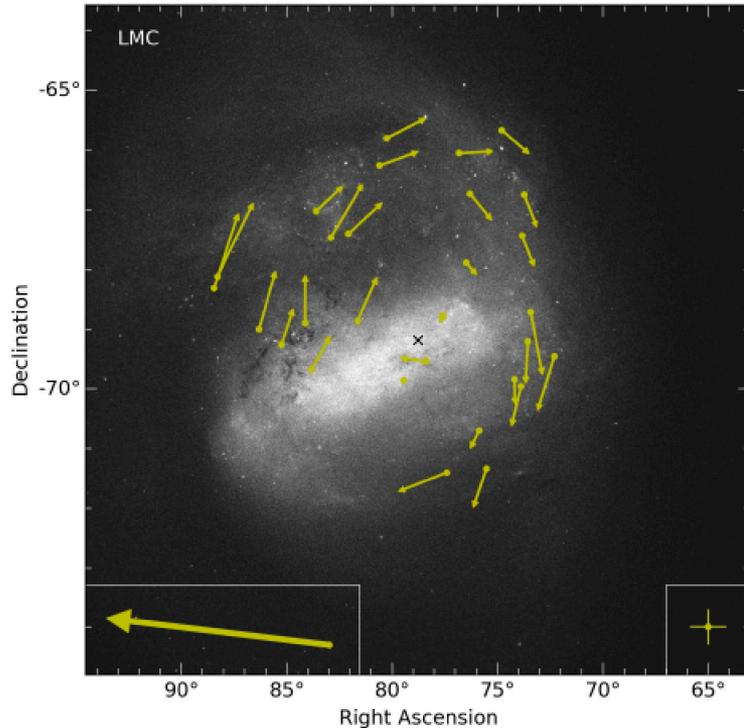
Hipparcos or Tycho-2 position (1991.25)

Starting with Gaia DR2, no prior information from Hipparcos and Tycho-2 will be used

→ Two distinct subsets in TGAS:

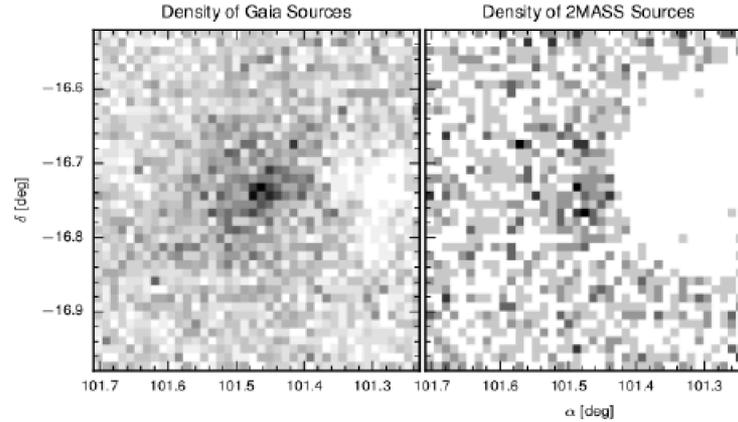
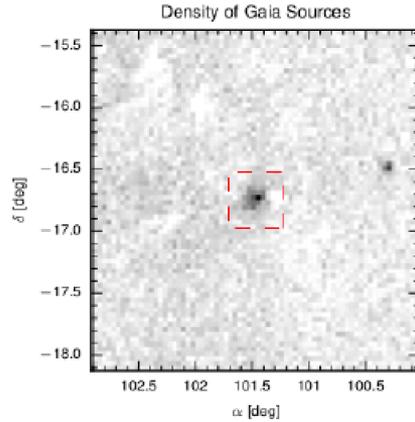
- Hip subset (~90,000 stars) using Hipparcos positions
- Tyc subset (~2 million stars) using Tycho-2 positions

- Parallax validation
  - ▶ No global offsets from Cepheids and RR Lyrae (PL relations)
  - ▶ Offsets claimed in samples of eclipsing binaries and in comparison to asteroseismic distances (Methodology? Assumptions about the stellar types involved?)
- Verification of quoted TGAS errors
  - ▶ Various indications of 10–20% overestimate of the errors: depends on sample of stars used
- Gaia DR1 positions and photometry already a reference for other surveys
  - ▶ new proper motion catalogues

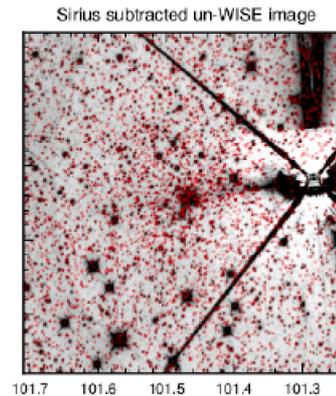
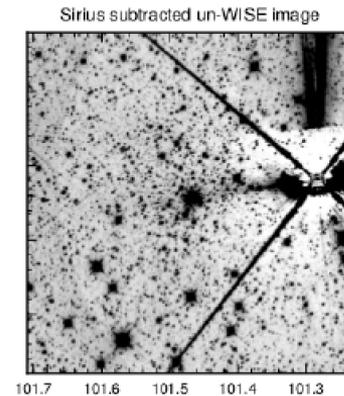
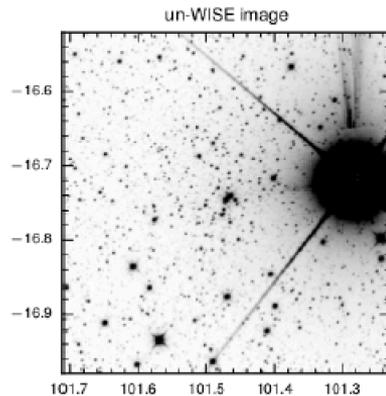


- Measured proper motions of LMC and SMC consistent with HST result
- Residual kinematics in LMC show clear rotation pattern
  - ▶ also seen in HST studies
  - ▶ rotation curve from proper motions consistent with result from line-of-sight motions
  - ▶ Kinematic distance modulus of  $18.54 \pm 0.39$

van der Marel & Sahlmann, arXiv:1609.04395

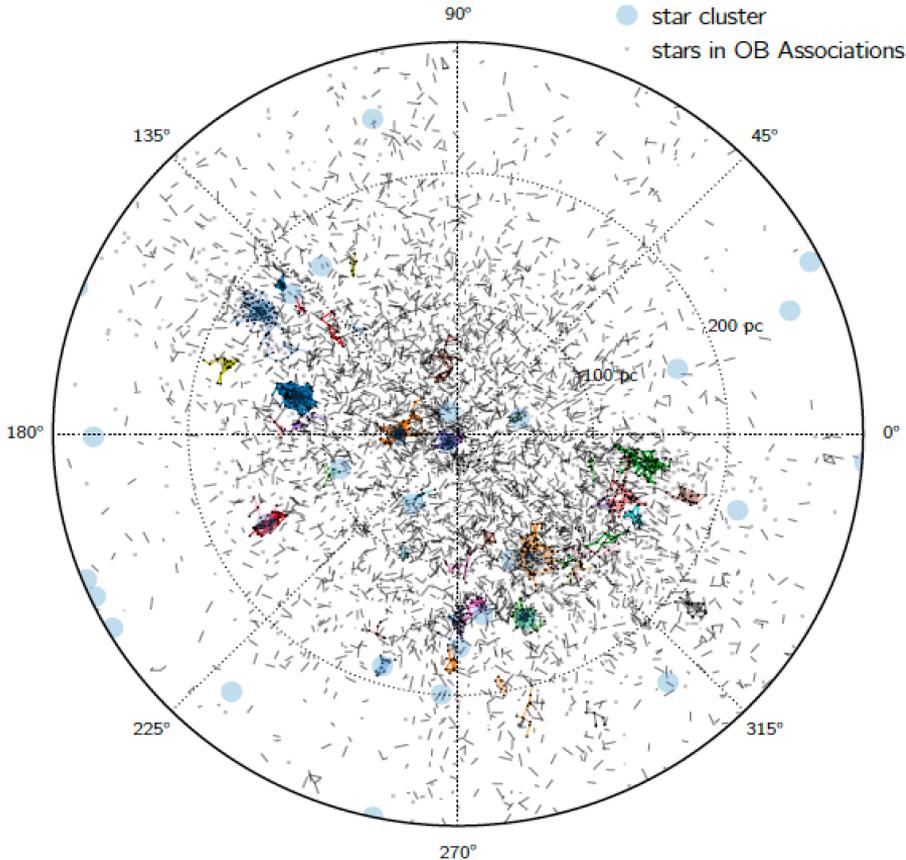


- Cluster hiding behind Sirius
- Power of all-sky, deep, high resolution, star-map



Koposov et al., 2017, arXiv:1702.01122

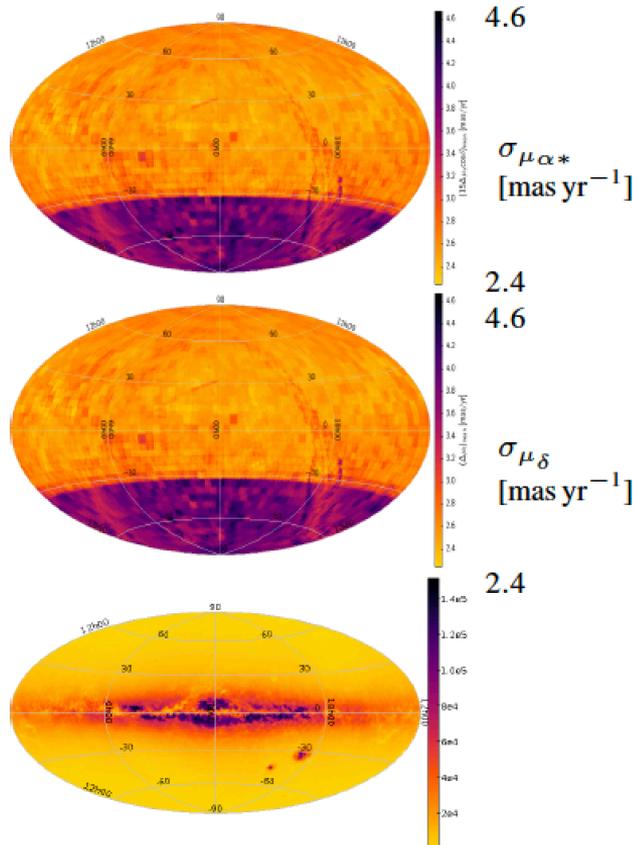
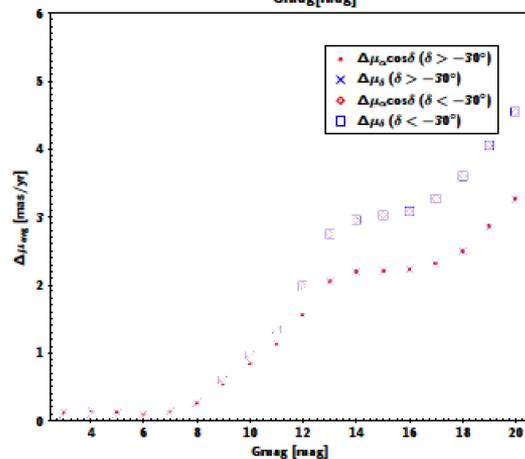
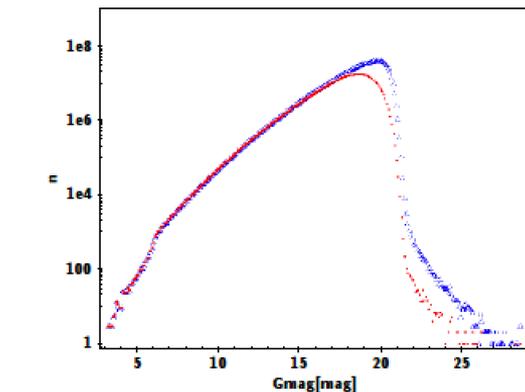
# Co-moving pairs



- 13 085 co-moving pairs identified in TGAS
  - ▶ based on proper motions, parallaxes, and marginalized model for 3D velocities
- Find pairs as well as networks of co-moving stars (clusters, associations, moving groups)
- Excess of pairs with  $> 1$  pc separation, wide binaries drifting apart?

Oh et al., arXiv:1612.02440

# Hot Stuff for One Year



- 583 million proper motions
- Gaia DR1 – PPMXL combination
- Altmann et al., 2017, A&A 600, L4 (arXiv:1701.02629)

- Five-parameter astrometric solutions for all sources with acceptable formal standard errors ( $> 10^9$  anticipated), and positions  $(\alpha, \delta)$  for sources for which parallaxes and proper motions cannot be derived
- $G$  and integrated  $G_{BP}$  and  $G_{RP}$  photometric fluxes and magnitudes for all sources
- Median radial velocities for sources brighter than  $G_{RVS} = 12$
- For stars brighter than  $G = 17$ , estimates of  $T_{\text{eff}}$  and, where possible,  $A_V$ , based on integrated photometry
- Photometric data for a sample of variable stars
- Epoch astrometry for a pre-selected list of  $> 10\,000$  asteroids